

3. Fermion fields

Clifford algebra : $\{\Gamma^M, \Gamma^N\} = 2\eta^{MN}$

$\Rightarrow \Sigma^{MN} \equiv \frac{i}{4} [\Gamma^M, \Gamma^N]$ satisfy Lorentz algebra (= "Spin $\frac{1}{2}$ rep." of L.A.)
; $d \geq 2$

• $d = 2R$ (even) : Γ^M are $2^R \times 2^R$ matrices, [in fundamental rep.]
constructed from Pauli σ^i [algorithm for $R \rightarrow R+1$]

• $\Gamma \equiv i^R \Gamma^0 \Gamma^1 \dots \Gamma^{2R-1} \equiv \gamma^5$ in 4D

\rightarrow eigenvalue ± 1 "chirality"

• $\{\Gamma, \Gamma^M\} = 0 \Rightarrow [\Gamma, \Sigma^{MN}] = 0$

\Rightarrow rep. is reducible (as rep. of Lorentz algebra)

$$\underline{2^R \text{ Dirac}} = \underline{2^{R-1} \text{ Weyl}} + \underline{2^{R-1} \text{ Weyl}}$$

\uparrow $\{ \psi | \Gamma \psi = \psi \}$ \uparrow $\{ \psi | \Gamma \psi = -\psi \}$

projection operator $P_{R,L} \equiv \frac{1}{2} (1 \pm \Gamma)$
 Γ in 4D : positive chirality
 $=$ helicity
 $=$ "right-handed"
 $\& \rightarrow -$, right \rightarrow left

 $\Rightarrow \psi = P_R \psi + P_L \psi \equiv \psi_R + \psi_L$

• $d = 2R+1$ (odd) : use $\{\Gamma^M\}$ for $d = 2R$ and add $\Gamma \equiv \Gamma^{2R}$

$\{\Gamma, \Gamma^M\} \neq 0$ (only for $M = 2R$!)

$\Rightarrow [\Gamma, \Sigma^{MN}] \neq 0$

\leadsto rep. is irreducible

consequences: chiral fermions

- must be massless (as in 4D, but valid for any d):

$$\mathcal{L} \supset m \bar{\psi} \psi = m(\bar{\psi}_R \psi_L + \bar{\psi}_L \psi_R)$$

i.e. a mass term would mix chiral states/
transform $L \leftrightarrow R$

- cannot exist in odd dimensions:

"L" and "R" are related by Lorentz trasfos!

Q: How to obtain chiral fermions in 4D?

[NB: even for $d \geq 6$, 5D Lorentz trasfos would mix eigenstates of $\frac{1 \pm \gamma^5}{2}$!]

A: compactify on an orbifold!

e.g. in 5D: $\psi = \psi_R + \psi_L$

do not mix under 5D Lorentz trasfos
4D = =

$\Rightarrow \psi_R, \psi_L$ can have different parities under $y \rightarrow -y$
(i.e. different properties under orbifold projections)

e.g. SM-like fermion "SU(2) singlet" (= "right-handed" in 4D)

$$\psi_5 = \frac{1}{\sqrt{2\pi R}} \psi_{5R}^{(0)} + \frac{1}{\sqrt{\pi R}} \sum_{n=1}^{\infty} \left(\psi_{5R}^{(n)} \cos \frac{n y}{R} + \psi_{5L}^{(n)} \sin \frac{n y}{R} \right)$$

bottom line: for every chiral fermion (in 4D) there are
two massive fermions at each KK level! (in a 5D theory)

4. Some historical notes

- Nordström (1914), Kaluza (1921) & Klein (1926)

→ consider gravity in 5D (in vacuum, no matter Lagrangian!):

$$S = \frac{1}{16\pi\hat{G}} \int d^5x \sqrt{|\hat{g}|} R^{(5)}$$

$\underbrace{R^{(5)}}_{= R^{\mu\nu}R_{\mu\nu}}$

$$= \frac{1}{16\pi\hat{G}} \int d^5x \sqrt{|\hat{g}|} \left\{ \underbrace{R^{\mu\nu}R_{\mu\nu}}_{R^{(4)}} + 2R^{\mu 5}R_{\mu 5} + R^{55}R_{55} \right\}$$



$$\bullet A_\mu \equiv (8\pi\hat{G})^{-1/2} \hat{g}_{5\mu}$$

+ assumptions: i) $y \sim y + 2\pi R$ n.b.: not the same "R" !!! Klein's "cylinder condition"

ii) $\partial_y \hat{g}_{\mu\nu} \approx 0$

(⇒ no KK tower (⇒ R very small)

iii) $\hat{g}_{55} = \text{const.}$ ← rather ad hoc!

[ii + iii ⇒ $R_{55} = 0$]

(...)

$$S = \int d^4x \left\{ \frac{1}{16\pi\hat{G}} R^{(4)} - \frac{1}{4} F^{\mu\nu} F_{\mu\nu} \right\} : \text{unification of GR \& EM!}$$

\uparrow
 $= \hat{G}/(2\pi R)$

- follow-up work:

- non-Abelian gauge groups possible for $d > 5$, depending on internal geometry

- $G_7 \supset SU(3) \times SU(2) \times U(1)$ possible for $d \geq 11$

- various issues...

quantizability (> 2 states for $d > 11$), renormalizability, chirality ..

- renewed interest with string theory (70's - 80's)



idea: finite extent \rightarrow cure UV (= short distance) problems

- consistent quantization possible only in $d=11$ (Minkowski)
 - \rightarrow discrete oscillation patterns: massless/light modes + KK towers
- open strings always attached to $d < 11$ objects ("branes")
 - \rightarrow various massless vector (gauge) fields in spectrum
- closed strings propagate freely in entire $d=11$ space ("bulk")
 - \rightarrow always contain massless spin-2 field \rightarrow gravity

- "large ED" scenario by Arkani-Hamed, Dimopoulos & Dvali (1998)

may explain hierarchy problem: (i.e. why is gravity so much weaker than other forces?)

- only gravity propagates in EDs
 - e.g. $d=6 + R \sim \text{mm}$
- $\Rightarrow M_{\text{pl}} = 10^{19} \text{ GeV} (\Rightarrow \hat{M}_{\text{pl}} \sim 10^2 \text{ GeV} \sim m_{\text{EW}})$

- Randall & Sundrum (1999)

- RS1: "warped" geometry allows solution of hierarchy also w/ small EDs

(anti-deSitter space in bulk: $ds^2 = e^{-2R|y|} (\sum_{\mu, \nu} dx^\mu dx^\nu + r_c^2 dy^2)$)

- RS2: Even infinitely large EDs can give Newton's law in 4D!
(but this no longer solves the hierarchy problem)

- ... [lots of (fading) activity; ADS+RS1/RS2 among most-cited theory papers of all times]

5. "Universal" extra dimensions

Can we put the entire SM in 5D?

Appelquist, Cheng & Dobrescu (2001): yes! If compactified on S^1/\mathbb{Z}_2 ...

- fermions: recall $G \supset SU(2)_L \leadsto$ doublets, e.g. $\begin{pmatrix} \nu_l \\ e_l \end{pmatrix}$
+ singlets, e.g. e_R

\rightarrow both singlets and doublets (=different chiralities in 4D)

come with two KK towers each

$\hat{=}$ right-/left-handed spinions in SUSY

- (massless) vectors - for $U(1), SU(2), SU(3)$ - come with each their KK tower

$\hat{=}$ gauginos

- scalars: Higgs doublet $\phi \sim \begin{pmatrix} \chi^0 + i\chi^1 \\ H - i\chi^2 \end{pmatrix}$ "cont+h" \checkmark give mass to W^\pm, Z in SM

$\leadsto \chi_i^{(n)}, \chi_i^{(n)}, \chi_3^{(n)}, h^{(n)}$ @ KK level

+ $A_5^{(n)}$: $1 \times$ for $U(1)$
 $3 \times$ = $SU(2)$ "W/Z"
 $8 \times$ = $SU(3)$ "gluons" $\left. \vphantom{\begin{matrix} 1 \times \\ 3 \times \\ 8 \times \end{matrix}} \right\}$ same quantum numbers as $\chi_i^{(n)}$
 \Rightarrow mix!

16 scalar states at each KK level

- 12 scalar d.o.f. ("foldstone modes") to give mass to $A_m^{(n)}$

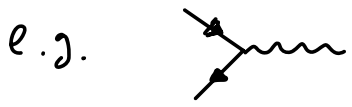
4 physical scalars for each KK level

$\hat{=}$ non-trivial Higgs sector in MSSM:

2 Higgs doublets \Rightarrow 5 physical scalars + 4 Higgsinos

still: UED sometimes referred to as "bosonic supersymmetry"

interactions

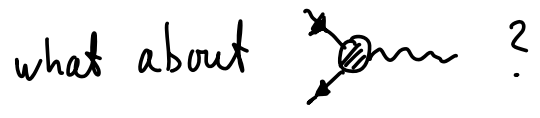


$$S \supset -\hat{e} \int d^4x dy A_\mu \bar{\psi}_s \gamma^\mu \psi_s$$

NB: Dirac, not Weyl, i.e. $\psi_s = \psi_{s,L} + \psi_{s,R}$

$$\approx -e \sum_{n, \ell, \ell'} \int d^4x A_\mu^{(n)} \bar{\psi}_s^{(\ell)} \gamma^\mu \psi_s^{(\ell')} \times \underbrace{\left\{ \delta_{n, \ell+\ell'} + \delta_{n, \ell-\ell'} + \delta_{n, \ell-\ell'} \right\}}_{\text{"KK-number conservation"}}$$

why? conservation of P_5 !



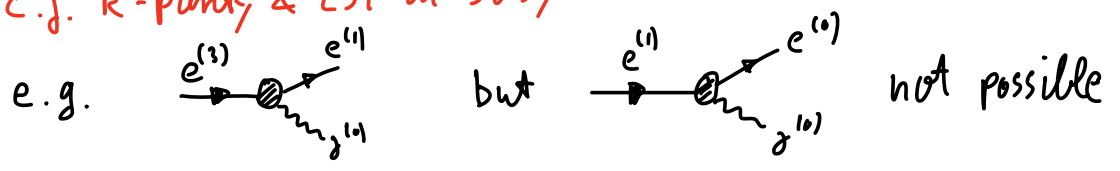
→ can not expect KK-number conservation!

why? P_5 not conserved @ $y=0, \pi R$



But $y \leftrightarrow -y \rightarrow$ still symmetry \Leftrightarrow "KK parity" $(-1)^n$ is conserved

\Rightarrow lightest KK particle (LKP) is absolutely stable

c.f. R-parity & LSP in SUSY



\Rightarrow spectrum of states phenomenologically very important

mass contributions: $\frac{1}{R}$, m_{EW} , , 

\uparrow same for all states
 \uparrow "tree-level" SM masses, from EW symmetry breaking
 \uparrow requires additional counterterms at orbifold boundaries \rightarrow dependence on cutoff scale Λ
 \uparrow finite + usual 4D infinities

result: $LKP = \delta^{(1)} \simeq \beta^{(1)}$ because Weinberg angle θ has at KK-0 level!

\rightarrow would be thermally produced in early universe
 \bullet = account for entire dark matter density (observed today) if $R^{-1} \sim 1.4 \text{ TeV}$

current constraint from LHC: $R^{-1} \gtrsim 1.7 \text{ TeV}$ [2110.00500]

\rightarrow † for KK DM...

† caveats apply: this is only for "minimal" VED.
 c.t. on orbifold fixpoints allow in principle for "arbitrary" $\delta m \Rightarrow$ required R^{-1} may be factor 2-3 higher